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Anisotropic Tilted Spherical False Vacuum Model

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ARTICLE INFO	ABSTRACT
	In this paper, a tilted spherically symmetric cosmological model with perfect fluid distribution
	along with heat conduction has been considered. To get a generalized solution, I have assumed
	the model to be False Vacuum (i.e. $p + \rho = 0$), where p being the isotropic pressure and
Corresponding Author: R.N.Patra	ho is the matter density of the fluid. The various physical properties of the model are also
	discussion as the concluding remark.
KEYWORDS: Tilted spherical cosmological model, False Vacuum model, Perfect fluid.	

1. INTRODUCTION

In the recent years, there has been a considerable interest in investigating spatially homogeneous and anisotropic cosmological models, in which matter do not move orthogonally to the hypersurface of homogeneity. These types of models are called tilted cosmological models. The general dynamics of these cosmological models have been studied in details by King and Ellis[1] and they have shown that in such universe, the matter move with non-zero expansion, rotation and shear. Also, Ellis and King[2], Collins and Ellis[3], Ellis and Baldwin[4] have shown that we are likely to be living in a tilted universe and they have indicated that how we may detect it.

The spherically symmetric cosmological models have been studied by several authors. S.N.Jena, R.N.Patra and R.R.Swain[5] have studied inhomogeneous macro model with perfect fluid, R.N.Patra[6,7] has studies in context of the cosmic string model, R.N.Patra et. al. [8, 9] have used spherical metric in presence of perfect fluid and magnetic field for construction of new metrices. U.K.Panigrahi and R.N.Patra and M. Sharma[10] have constructed metric by using new scalar tensor theory of Gravitation.

Now I am motivated by these studies and taken interest to construct a tilted spherically symmetric cosmological model filled with perfect fluid along with heat conduction.

2. THE FIELD EQUATION

We consider the spherically symmetric space-time

$$ds^{2} = -e^{\lambda}dr^{2} - r^{2}d\theta^{2} - r^{2}\sin^{2}\theta \,d\phi^{2} + e^{\mu}dt^{2},$$
(1)

where $\lambda(t)$ and $\mu(t)$ are the cosmic parameter.

The Energy momentum tensor for perfect fluid distribution with heat conduction given by Ellis[10] is given by

$$T_{i}^{j} = (p + \rho) v_{i} v^{j} + P g_{i}^{j} + q_{i} v^{i} + v_{i} q^{j}$$
(2)

together with

(4)

$$\begin{cases}
g_{ij} v^{i} v_{j} = -1 \\
q^{i} q^{j} > 0 \\
q_{i} v_{i} = 0
\end{cases}$$
(3)

Here, p is the isotropic pressure, ρ is the matter density, q_i the heat conduction vector orthogonal to v_i . The fluid flow vector v_i has the components

 $(\sin h\alpha, 0, 0, \cosh \alpha)$ satisfying equation (3), where α being the tilt angle.

The Einstein's field equation

$$R_i^j - \frac{1}{2} Rg_i^j = -8\pi T_i^j, (C = G = 1).$$

The field equation for the line element (1) leads to $-\frac{1}{4}\lambda_4^2 e^{\lambda-\mu} - \frac{\lambda_4 \mu_4}{8} e^{\lambda-\mu} = -8\pi \left\{ (p+\rho) \sin h\alpha - e^{\lambda}p + 2q_1 \sin h\alpha \right\}$

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$$-\frac{\lambda_4}{r} = -8\pi \left\{ \left(p + \rho \right) \sin h\alpha \cos h\alpha + q_1 \cos h\alpha + v_1 \frac{\sin h^2 \alpha}{\cos h\alpha} \right\}$$
(5)

$$\frac{\lambda_{44}}{2} + \frac{\lambda_4^2}{2} - \frac{\lambda_4 \mu_4}{8} = 8\pi \ p e^{\mu} \tag{6}$$

$$-\frac{1}{4}\lambda_4^2 + \frac{1}{r^2}e^{\mu-\lambda} - \frac{1}{r^2}e^{\mu} = -8\pi\{(p+\rho)\cos h^2\alpha + e^{\mu}p + 2q_1\sin h\alpha\}$$
(7)

3. SOLUTION TO THE FIELD EQUATIONS

Taking False Vacuum model $(\rho = -p)$, the tilt angle $\alpha = 1$, and the approximate value of e = 2, the value of λ from equation (5) is found to be

$$\lambda = 13.6rt q_1$$
.

Choosing the value of the heat conduction vector $q_1 = \frac{1}{13.6} = 0.074$, the value of λ becomes

$$\lambda = rt. \tag{8}$$

Substracting equation (6) multiplied by $e^{\lambda-\mu}$ from equation (4) and using the above values of ρ, α, λ, e and q_1 , the value of μ found to be

$$\mu = 3r^3t - 2.4$$
. (9)

The heat conduction vectors q_1 and q_4 are found to be in a relation

$$q_4 = 3q_1$$
 (10)
 $q_1 = 0.074$, so $q_4 = 0.222$.

From equation (6), the value of the isotropic pressure is found to be

$$p = \frac{4r^2 - 3r^4}{201e^{3r^3t - 2.4}}.$$
 (11)

Using the values of λ and μ , the new form of the space-time

becomes

$$ds^{2} = -e^{rt} dr^{2} - r^{2} d\theta^{2} - r^{2} \sin^{2} \theta d\phi^{2} + e^{3r^{3}t - 2.4} dt^{2}.$$

4. PHYSICAL PROPERTIES

1. Assuming the heat conduction vector along the fourth co-ordinate to be thrice of that along the first co-ordinate, the isotropic pressure and the matter density of the space-time with perfect fluid distribution as in equation (11) is

$$p = -\rho = \frac{4r^2 - 3r^4}{201\,e^{3r^3t - 2.4}}$$

When $t - \infty$ then $\rho \to 0$ with a negative pressure. The variation of ρ w.r.t 't' has been shown in the graph, where ρ is a decreasing functions of time.



2. Taking the value of the tilt angle $\alpha = 1$ and the approximate value of e = 2, the fluid flow and the heat conduction vectors are found to be

$$v_1 = \sin h\alpha = 0.75$$

 $v_4 = \cos h\alpha = 1.25$ $q_1 = 0.074 \text{ and } q_4 = 0.222$.

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As the tilt angle is independent of time, the velocity components are found to be constant, that means the flow is uniform.

$$H = \frac{2r^{3}t + 3r^{4} - 2.4}{2r(3r^{3}t - 2.4)}$$
 is a function of

3. The Hubble parameter

'r' and 't'.
The
$$H \sim t$$
 graph has been shown below.





 $V \rightarrow \infty as t \rightarrow \infty$ and $V \rightarrow 0 as t \rightarrow 0$ that means the universe expands with the increase of cosmic time and may blow up in near future which has been shown in the graph.



5.



found to be

i.e. q < 0, which shows the accelerating nature of the universe. The graph of deceleration parameter q has been

shown in the graph given below.



6. The mean anisotropy parameter

$$A = \frac{1}{2} \left(\frac{\Delta H_1}{H} + \frac{\Delta H_2}{H} \right) \neq 0$$

That means the model of the universe is found to be anisotropic.

5. CONCLUSION

The false vacuum model $(p = -\rho)$ applied to the universe filled with perfect fluid and heat conduction is found to be anisotropic with expansion in volume as the deceleration parameter q < 0. Taking the tilt angle $\alpha = 1$, the flow of the fluid is found to be uniform. The matter density of the model decreases sharply with increase in cosmic time and the isotropic pressure becomes less negative gradually. Taking the approximate value of e = 2, the heat conduction vectors are found to be in a relation $q_4 = 3q_1$.

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